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Higgs phenomenology in the Peccei-Quinn invariant NMSSM

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ABSTRACT: We study the Higgs phenomenology in the Peccei-Quinn invariant NMSSM (PQ-NMSSM) where the low energy mass parameters of the singlet superfield are induced by a spontaneous breakdown of the Peccei-Quinn symmetry. In the generic NMSSM, scalar mixing among CP-even Higgs bosons is constrained by the observed properties of the SM-like Higgs boson, as well as by the LEP bound on the chargino mass and the perturbativity bound on the singlet Yukawa coupling. In the minimal PQ-NMSSM, scalar mixing is further constrained due to the presence of a light singlino-like neutralino. It is noticed that the 2σ excess of the LEP $Zb\bar{b}$ events at $m_{b\bar{b}} \simeq 98$ GeV can be explained by a singlet-like 98 GeV Higgs boson in the minimal PQ-NMSSM with low $\tan\beta$, stops around or below 1 TeV, and light doublet-higgsinos around the weak scale.

KEYWORDS: Supersymmetry Phenomenology

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1 Introduction

There are many reasons to anticipate new physics beyond the standard model (SM), including the naturalness problems such as the hierarchy problem and the strong CP problem, and a variety of cosmological observations such as the existence of dark matter, the matter-antimatter asymmetry, and the evidences for inflation in the early Universe. Among the known scenarios of new physics, a particularly compelling possibility is a supersymmetric extension of the SM [1, 2] incorporating also the axion solution to the strong CP problem through a spontaneously broken Peccei-Quinn (PQ) symmetry [3–5]. While solving the two major naturalness problems of the SM, such an extension of the SM provides an attractive candidate for dark matter, either the lightest supersymmetric particle or the axion, or both. It also offers an interesting possibility that the PQ scale is generated by an interplay between supersymmetry (SUSY) breaking effect and Planck-scale suppressed effect, yielding an intermediate PQ scale $v_{\text{PQ}} \sim \sqrt{m_{\text{soft}} M_{\text{Pl}}}$ in a natural manner [6, 7], where m_{soft} is a soft SUSY breaking mass presumed to be of the order of the weak scale. In such a scenario, the PQ phase transition takes place in the early Universe at a temperature $T \sim m_{\text{soft}}$. This results in a late thermal inflation over the period $m_{\text{soft}} < T < v_{\text{PQ}}$, with which dangerous cosmological relics such as the moduli and gravitinos are all diluted away [7–11].

The scalar boson with a mass $m_h \simeq 125$ GeV, which was recently discovered in the LHC experiments, has been found to behave like the SM Higgs boson [12, 13]. On the other hand, a SM-like Higgs boson at 125 GeV in the minimal supersymmetric standard model (MSSM) requires that stops have either a heavy mass in multi-TeV range or maximal LR-mixing,

which would cause a fine-tuning worse than 1 % in the electroweak symmetry breaking. This fine-tuning can be ameliorated in the next-to-minimal supersymmetric standard model (NMSSM) involving a singlet superfield S with the superpotential coupling $\lambda SH_u H_d$. In the NMSSM, the SM-like Higgs boson h gains an additional tree-level mass from the F -term scalar potential $\lambda^2 |H_u H_d|^2$, or from scalar mixing if the singlet scalar s is lighter than h . This makes it possible to have $m_h \simeq 125 \text{ GeV}$ even when stops are relatively light and stop mixings are small, and therefore reduces the amount of fine-tuning required for the electroweak symmetry breaking [14–19]. Furthermore the Higgs and neutralino sector of the NMSSM have a richer structure than the MSSM. If all the Higgs bosons in the NMSSM have masses in sub-TeV range, e.g. below 500 GeV, there can be sizable mixings among the three CP-even Higgs bosons, leading to interesting phenomenological consequences as discussed in [16–28].

It is well known that a PQ-symmetry spontaneously broken at $v_{\text{PQ}} \sim \sqrt{m_{\text{soft}} M_{\text{Pl}}}$ can explain why the doublet-higgsino mass μ in the MSSM is comparable to m_{soft} [6, 7, 29–31]. Similarly, if the singlet superfield S is PQ-charged, the low energy mass parameters of S in the effective superpotential of the PQ-invariant NMSSM are induced by a spontaneous breakdown of the PQ symmetry, and so can have a value comparable to m_{soft} , while the singlet cubic coupling is always negligible [32–34]. In this paper, we wish to examine the Higgs phenomenology in such a PQ-invariant NMSSM while focusing on the phenomenological consequences of scalar mixing. The Higgs boson masses and mixing angles in the neutral CP-even Higgs sector crucially depend on the coupling λ and the doublet-higgsino mass μ . As a result, scalar mixing is constrained not only by the observed mass and signal strengths of the SM-like Higgs boson, but also by the perturbativity bound on λ and the LEP bound on the chargino mass. We will examine first the constraints on scalar mixing in the context of the general NMSSM, and then consider a specific minimal PQ-invariant NMSSM which is further constrained by the presence of a light singlino-like neutralino.

If the singlet-like Higgs boson s has a mass near the weak scale, it can have a large mixing with the SM-like Higgs boson h . We identify the parameter region of the sizable singlet-doublet mixing that is compatible with all the LHC and LEP data available at present, as well as with the perturbativity bound on λ and a stop mass between 600 GeV and a few TeV. We explore also the possibility that the 2σ excess of the LEP $Zb\bar{b}$ events at $m_{b\bar{b}} \simeq 98 \text{ GeV}$ is explained by $e^+e^- \rightarrow Zs \rightarrow Zb\bar{b}$ within the framework of the minimal PQ-NMSSM.¹ We then find that it requires low $\tan\beta$ smaller than about 2, a light doublet-higgsino mass around the weak scale, and stop masses around or below 1 TeV. For the case with $m_s > m_h$, it is found that s decays dominantly into a neutralino pair in most of the viable parameter region, which would make its detection at collider experiments difficult. We examine also the signal rates of the SM-like Higgs boson in the $b\bar{b}$ ($\tau\bar{\tau}$) and di-photon channels over the phenomenologically viable parameter region which gives the signal rate of the WW/ZZ channel close to the SM value.

This paper is organized as follows. In section 2, we discuss how the constraints on λ and μ , and the observed properties of the SM-like Higgs boson translate into the constraints on

¹Such a possibility for the conventional Z_3 -invariant NMSSM has been examined recently in ref. [35].

the scalar mixing angles. In section 3, we discuss some generic features of the PQ-invariant NMSSM, and present a specific model which is considered to be a minimal PQ-invariant NMSSM with $v_{\text{PQ}} \sim \sqrt{m_{\text{soft}} M_{\text{Pl}}}$. The neutralino sector of the minimal PQ-NMSSM is also discussed with a focus on the additional constraints arising due to a light singlino-like neutralino in the model. In section 4, we apply the results of the section 2 to the Higgs phenomenology in the minimal PQ-NMSSM. We present first the results that hold in the general NMSSM, and then impose additional constraints specific to the minimal PQ-NMSSM. Section 5 is the conclusions.

2 Constraints on Higgs mixing in the NMSSM

In this section, we briefly discuss phenomenological consequences of Higgs mixing in the general NMSSM and the resultant constraints on the model. Let us begin with the Higgs sector superpotential of the general NMSSM, which is given by

$$\lambda S H_u H_d + f(S), \quad (2.1)$$

in an appropriate basis of the singlet superfield S . The first term is responsible for the higgsino mass parameter μ and Higgs bilinear coupling $B\mu$:

$$\begin{aligned} \mu &= \lambda \langle S \rangle, \\ B\mu &= \lambda (A_\lambda \langle S \rangle + \langle \partial_S f \rangle), \end{aligned} \quad (2.2)$$

where A_λ is the soft SUSY breaking parameter for the superpotential term $S H_u H_d$. There is one combination \hat{h} of CP-even neutral Higgs bosons which corresponds to the fluctuation of $\text{Re}(H_u^0)$ and $\text{Re}(H_d^0)$ in the vacuum value direction, and therefore behaves like the SM Higgs boson in the limit when the other Higgs bosons are decoupled. In the NMSSM, it generally mixes with the other CP-even neutral Higgs bosons, and the SM-like Higgs boson in the mass-eigenstate is given by

$$h = c_{\theta_1} c_{\theta_2} \hat{h} - s_{\theta_1} \hat{H} - c_{\theta_1} s_{\theta_2} \hat{s}, \quad (2.3)$$

with $c_{\theta_i} = \cos \theta_i$ and $s_{\theta_i} = \sin \theta_i$ for the mixing angles θ_i defined in appendix A, where \hat{H} is the fluctuation of $\text{Re}(H_u^0)$ and $\text{Re}(H_d^0)$ orthogonal to \hat{h} , and \hat{s} is the CP-even fluctuation of the singlet scalar.

Around the weak scale, the SM-like Higgs boson interacts with the SM particles through the terms² [36],

$$\begin{aligned} \mathcal{L} &= C_V \frac{\sqrt{2} m_W^2}{v} h W_\mu^+ W_\mu^- + C_V \frac{m_Z^2}{\sqrt{2} v} h Z_\mu Z_\mu - C_f \frac{m_\psi}{\sqrt{2} v} h \bar{f} f \\ &\quad + C_g \frac{\alpha_s}{12 \sqrt{2} \pi v} h G_{\mu\nu}^a G_{\mu\nu}^a + C_\gamma \frac{\alpha}{\sqrt{2} \pi v} h A_{\mu\nu} A_{\mu\nu}, \end{aligned} \quad (2.4)$$

²This should be understood as an 1PI effective Lagrangian including quantum corrections for the SM-like Higgs boson near the mass-shell.

where f denote the SM fermions, and $v \simeq 174 \text{ GeV}$ is the Higgs vacuum expectation value. At tree level, the Higgs couplings to the massive SM particles are determined by the mixing angles as

$$\begin{aligned} C_V &= c_{\theta_1} c_{\theta_2}, \\ C_t &= c_{\theta_1} c_{\theta_2} + s_{\theta_1} \cot \beta, \\ C_b &= C_\tau = c_{\theta_1} c_{\theta_2} - s_{\theta_1} \tan \beta. \end{aligned} \quad (2.5)$$

On the other hand, the Higgs couplings to massless gluons and photons are radiatively induced. The dominant contribution comes from the W -boson and top quark loops:

$$\begin{aligned} C_g &\simeq 1.03 C_t - 0.06 C_b + \delta C_g, \\ C_\gamma &\simeq 0.23 C_t - 1.04 C_V + \delta C_\gamma, \end{aligned} \quad (2.6)$$

including superparticle loop contributions δC_g and δC_γ , where δC_g can be sizable if stops are below 1 TeV, and δC_γ becomes important if there are light charged-superparticles around the weak scale. Using the above relations, one can estimate the signal rate of the SM-like Higgs boson h at the LHC in the presence of scalar mixing. The signal rate in the WW/ZZ channel normalized by the SM value is given by

$$R_h^{VV} \simeq \frac{(0.94 C_g^2 + 0.12 C_V^2) C_V^2}{0.64 C_b^2 + 0.24 C_V^2 + 0.12 C_t^2}, \quad (2.7)$$

where we have used the well-known production and decay properties of the SM Higgs boson under the assumption that the Higgs decay rate into non-SM particles is negligible. To see the effect of scalar mixing, it is convenient to factor the signal rate into WW/ZZ as

$$R_h^{VV} \simeq \left(1 + 2 \frac{\delta C_g}{C_g}\right) R_h^{VV}|_0, \quad (2.8)$$

where $R_h^{VV}|_0$ is the signal rate for $\delta C_g = 0$, i.e. in the limit that all the colored superparticles are heavy. It is important to note that $R_h^{VV}|_0$ depends only on θ_1 , θ_2 , and $\tan \beta$. In addition, because the effect of colored superparticles is to modify the Higgs production rate in the gluon fusion process, the ratio R_h^{ii}/R_h^{VV} for each channel is insensitive to the correction δC_g . For other channels, we find

$$\begin{aligned} R_h^{bb} &= R_h^{\tau\tau} = \frac{C_b^2}{C_V^2} R_h^{VV}, \\ R_h^{\gamma\gamma} &\simeq \frac{1.52 C_\gamma^2}{C_V^2} R_h^{VV}, \end{aligned} \quad (2.9)$$

where $R_h^{ii} = 1$ in the limit of vanishing mixing angles and decoupled superparticles.

In the NMSSM, the Higgs quartic coupling receives an additional tree-level contribution proportional to λ^2 , and consequently \hat{h} obtains a mass according to

$$m_{\hat{h}}^2 = m_0^2 + (\lambda^2 v^2 - m_Z^2) \sin^2 2\beta, \quad (2.10)$$

where m_0 corresponds to the SM-like Higgs boson mass at large $\tan \beta$ in the decoupling limit of MSSM, including the well-known radiative correction from top and stop loops [37–39]:

$$m_0^2 = m_Z^2 + \frac{3m_t^4}{4\pi^2 v^2} \ln \left(\frac{m_{\tilde{t}}^2}{m_t^2} \right) + \frac{3m_t^4}{4\pi^2 v^2} \left(X_t^2 - \frac{1}{12} X_t^4 \right) + \dots, \quad (2.11)$$

for the stop mass $m_{\tilde{t}}$ and the stop mixing parameter $X_t = (A_t - \mu \cot \beta)/m_{\tilde{t}}$. It is straightforward to see that the mass of the SM-like Higgs boson in the NMSSM reads

$$m_h^2 = m_{\tilde{h}}^2 - \frac{(s_{\theta_2} s_{\theta_3} - s_{\theta_1} c_{\theta_2} c_{\theta_3})^2}{c_{\theta_1}^2 c_{\theta_2}^2} (m_H^2 - m_{\tilde{h}}^2) - \frac{(s_{\theta_2} c_{\theta_3} + s_{\theta_1} c_{\theta_2} s_{\theta_3})^2}{c_{\theta_1}^2 c_{\theta_2}^2} (m_s^2 - m_{\tilde{h}}^2), \quad (2.12)$$

where the last two terms are due to scalar mixing. Note that the mixing with singlet scalar increases m_h^2 if the singlet-like Higgs boson s is lighter than the SM-like Higgs boson h [32, 33, 40, 41].

In the presence of scalar mixing, the singlet-like Higgs boson s also interacts with the SM particles via the doublet components. Those interactions are obtained from (2.4) by replacing C_i with the effective couplings

$$\begin{aligned} C_V^s &= s_{\theta_2} c_{\theta_3} + s_{\theta_1} c_{\theta_2} s_{\theta_3}, \\ C_t^s &= s_{\theta_2} c_{\theta_3} + s_{\theta_1} c_{\theta_2} s_{\theta_3} - c_{\theta_1} s_{\theta_3} \cot \beta, \\ C_b^s &= C_\tau^s = s_{\theta_2} c_{\theta_3} + s_{\theta_1} c_{\theta_2} s_{\theta_3} + c_{\theta_1} s_{\theta_3} \tan \beta, \end{aligned} \quad (2.13)$$

at tree-level, and the coupling to gluons and photons are radiatively generated depending on the singlet mass m_s .

Let us examine how the SM-like Higgs boson in the NMSSM can be arranged to be consistent with the LHC data. The most important constraints come from the mass and signal rates for the various Higgs decay channels observed at the LHC. In particular, the signal rate for the WW/ZZ channel should be close to the SM value,

$$R_h^{VV} \approx 1, \quad (2.14)$$

which does not necessarily imply that the h - s mixing angle θ_2 should be small. Keeping in mind that the Higgs coupling to gluons can receive a non-negligible correction from relatively light stops, we impose the condition

$$R_h^{VV}|_0 \simeq 1, \quad (2.15)$$

to account for the observed Higgs signal rate in WW/ZZ . This is the case when the mixing angles obey the relation [25],

$$\theta_1 \approx \frac{\tan \beta}{1.4 \tan^2 \beta + 1.7} \sin^2 \theta_2. \quad (2.16)$$

Here we have used that $R_h^{VV}|_0$ is determined only by θ_1 , θ_2 , and $\tan \beta$. For such Higgs mixing, the signal rates for the fermionic ($b\bar{b}$ or $\tau\bar{\tau}$) and di-photon channel are estimated to be

$$\begin{aligned} R_h^{bb} &= R_h^{\tau\tau} \approx (1 - \theta_1 \tan \beta)^2 R_h^{VV}, \\ R_h^{\gamma\gamma} &\approx (1 - 0.28\theta_1 \cot \beta - 1.23 \delta C_\gamma)^2 R_h^{VV}, \end{aligned} \quad (2.17)$$

with $R_h^{VV} \approx 1$. This shows that the signal rates for the $b\bar{b}$ and $\tau\bar{\tau}$ channel are reduced below the SM prediction as a result of scalar mixing at tree level. The di-photon rate is less affected by scalar mixing. However, in the presence of sizable θ_2 and light charged-higgsinos, it can significantly deviate from the SM value due to the chargino-loop contribution to δC_γ , which is given by [36]

$$\delta C_\gamma|_{\tilde{H}^\pm} \approx -0.17 \frac{\lambda v}{|\mu|} \tan \theta_2. \quad (2.18)$$

Note that the charged-higgsino loop can either enhance or reduce the di-photon rate, depending on the sign of θ_2 .

In the NMSSM, for a given value of $\tan \beta$, the off-diagonal components of the mass matrix of $(\hat{h}, \hat{H}, \hat{s})$ are determined by three parameters $\{\lambda, \mu, \Lambda\}$ (see appendix A), where

$$\Lambda = A_\lambda + \langle \partial_S^2 f \rangle \quad (2.19)$$

is independent from the effective Higgs bilinear coupling $B\mu$. These parameters can be expressed in terms of the mixing angles θ_i and the mass eigenvalues m_h, m_H and m_s . In particular, λ and μ are given by

$$\begin{aligned} \lambda^2 v^2 &= m_Z^2 + \frac{1}{\sin 4\beta} \left((m_H^2 - m_s^2) s_{\theta_2} s_{2\theta_3} + 2(m_h^2 - m_H^2 c_{\theta_3}^2 - m_s^2 s_{\theta_3}^2) s_{\theta_1} c_{\theta_2} \right) c_{\theta_1}, \\ \lambda v \mu &= -\frac{1}{4} m_h^2 c_{\theta_1}^2 s_{2\theta_2} - \frac{1}{4} (m_H^2 - m_s^2) s_{\theta_1} c_{2\theta_2} s_{2\theta_3} \\ &\quad + \frac{1}{4} \left((m_H^2 - m_s^2 s_{\theta_1}^2) s_{\theta_3}^2 - (m_H^2 s_{\theta_1}^2 - m_s^2) c_{\theta_3}^2 \right) s_{2\theta_2} \\ &\quad - \frac{\tan 2\beta}{4} \left((m_H^2 - m_s^2) c_{\theta_2} s_{2\theta_3} - 2(m_h^2 - m_H^2 c_{\theta_3}^2 - m_s^2 s_{\theta_3}^2) s_{\theta_1} s_{\theta_2} \right) c_{\theta_1}. \end{aligned} \quad (2.20)$$

On the other hand, the coupling λ is constrained to be less than about 0.7 at the weak scale, if one wishes to maintain the model to be perturbative up to the GUT scale [42], while the LEP bound on the chargino mass requires $|\mu|$ to be larger than about 100 GeV [43]. These constraints on λ and μ can be translated into those on the mixing angles and mass eigenvalues through the above relations.

Finally, m_0 cannot take an arbitrary value, and thus Higgs mixing is constrained by the requirement $m_h \simeq 125$ GeV through the relation (2.12). For instance, the stop searches at the LHC suggest that the stop is heavier than about 600 GeV [44–46], implying $m_0 \gtrsim 105$ GeV. One may avoid this stop mass bound by considering the case where the stop mass is smaller than the sum of the top quark mass and the lightest neutralino mass. On the other hand, fine-tuning for the electroweak symmetry breaking becomes more severe for heavier stop masses, so the naturalness principle favors m_0 to be as small as possible. We therefore assume m_0 to be in the range,

$$105 \text{ GeV} \lesssim m_0 \lesssim 120 \text{ GeV}, \quad (2.21)$$

which amounts to assuming that stops are not significantly heavier than 1 TeV. Note that δC_g receives the dominant contribution from stop loops [47, 48],

$$\delta C_g \approx \frac{1}{4} \frac{m_t^2}{m_t^2} (2 - X_t^2) C_t, \quad (2.22)$$

which can be sizable for stop masses of our interest. This correction to the Higgs coupling to gluons modifies the Higgs production rate in the gluon fusion, and enhances the Higgs signal rates for the stop mixing parameter $X_t < \sqrt{2}$. For instance, taking $X_t = 0$, one finds that the Higgs signal rate in each channel increases universally by about 8 % for the stop mass around 600 GeV, and by less than about 3 % for the stop mass heavier than 1 TeV. Here the ratio R_h^{ii}/R_h^{VV} remains almost the same. The stop contribution to the Higgs-photon coupling C_γ is below 1 % even for the stop mass around 600 GeV.

We close this section by summarizing the conditions yielding constraints on the Higgs mixing in the general NMSSM. These include (a) the mass of the SM-like Higgs boson, $m_h \simeq 125$ GeV, (b) the Higgs signal rates, in particular $R_h^{VV} \approx 1$, (c) the perturbativity bound $\lambda \lesssim 0.7$, and (d) the LEP bound on the chargino mass, implying $|\mu| \gtrsim 100$ GeV.

3 Peccei-Quinn invariant NMSSM

In this section we discuss the generic low energy limit of the PQ-invariant NMSSM, and present a specific model considered to be a minimal PQ-NMSSM. As we will see, a key feature of the minimal PQ-NMSSM is the presence of a light singlino-like neutralino, with which the model is severely constrained by the Higgs invisible decay and the LEP bound on neutralino productions.

3.1 Low energy limit of the generic PQ-NMSSM

At energy scales below the PQ-breaking scale v_{PQ} , the PQ-NMSSM can be described by a low energy effective theory with a non-linear $U(1)_{\text{PQ}}$ symmetry, under which the NMSSM Higgs superfields and the axion superfield A transform as

$$\begin{aligned} S &\rightarrow e^{i\alpha} S, \\ H_u H_d &\rightarrow e^{-i\alpha} H_u H_d, \\ A &\rightarrow A + i v_{\text{PQ}} \alpha. \end{aligned} \tag{3.1}$$

Throughout this paper, we assume that the PQ-breaking scale is generated by competition between SUSY breaking effect and Planck scale suppressed effect, so that

$$v_{\text{PQ}} \sim \sqrt{m_{\text{soft}} M_{\text{Pl}}}.$$

Here the axion superfield A is composed of a pseudo-scalar axion a solving the strong CP problem, its scalar partner saxion ρ , and the fermionic partner axino \tilde{a} :

$$A = \frac{1}{\sqrt{2}}(\rho + ia) + \sqrt{2}\theta\tilde{a} + \theta^2 F^A. \tag{3.2}$$

The PQ-invariant Kähler potential and superpotential below v_{PQ} are generically given by

$$\begin{aligned} K &= K_0(A + A^*) + \sum_i Z_i(A + A^*)|\Phi_i|^2 + \Delta K, \\ W &= (\text{MSSM Yukawa terms}) + \lambda S H_u H_d + \Delta W, \end{aligned} \tag{3.3}$$

in which Φ_i denote the NMSSM chiral superfields. Here ΔK and ΔW stand for the terms induced by a spontaneous breakdown of the PQ symmetry,

$$\begin{aligned}\Delta K &= \tilde{\mu}_1 e^{A^*/v_{\text{PQ}}} S + \kappa_1 e^{2A^*/v_{\text{PQ}}} S^2 + \kappa_2 e^{-A^*/v_{\text{PQ}}} H_u H_d + \dots + \text{h.c.}, \\ \Delta W &= \tilde{\mu}_2^2 e^{-A/v_{\text{PQ}}} S + \tilde{\mu}_3 e^{-2A/v_{\text{PQ}}} S^2 + \tilde{\mu}_4 e^{A/v_{\text{PQ}}} H_u H_d + \kappa_3 e^{-3A/v_{\text{PQ}}} S^3 + \dots, \end{aligned} \quad (3.4)$$

where the ellipses denote higher dimensional terms, and

$$\kappa_j \lesssim \mathcal{O}\left((v_{\text{PQ}}/M_{\text{Pl}})^{k_j}\right), \quad \frac{\tilde{\mu}_j}{v_{\text{PQ}}} \lesssim \mathcal{O}\left((v_{\text{PQ}}/M_{\text{Pl}})^{n_j}\right), \quad (3.5)$$

for model-dependent non-negative integers k_j and n_j .

Including the effects of soft SUSY-breaking, the vacuum value of the axion superfield can be determined to be³

$$\frac{\langle A \rangle}{v_{\text{PQ}}} = \xi_1 + \xi_2 m_{\text{soft}} \theta^2, \quad (3.6)$$

where $\xi_{1,2} = \mathcal{O}(1)$ in general. To examine the particle physics phenomenology at scales below v_{PQ} , it is convenient to replace the axion superfield with its vacuum expectation value. After this replacement, one can make an appropriate field redefinition

$$S \rightarrow S + \mu_0 + b_0 \theta^2, \quad (3.7)$$

together with a Kähler transformation

$$K_{\text{eff}} \rightarrow K_{\text{eff}} - (\bar{\theta}^2 \Omega + \text{h.c.}), \quad W_{\text{eff}} \rightarrow W_{\text{eff}} + \Omega, \quad (3.8)$$

to arrive at the following form of the effective Kähler potential and superpotential:

$$\begin{aligned} K_{\text{eff}} &= \sum_i (1 - m_i^2 \theta^2 \bar{\theta}^2) |\Phi_i|^2, \\ W_{\text{eff}} &= (\text{MSSM Yukawa and } A\text{-terms}) + \lambda(1 + A_\lambda \theta^2) S H_u H_d \\ &\quad + \mu_1^2 (1 + B_1 \theta^2) S + \frac{1}{2} \mu_2 (1 + B_2 \theta^2) S^2 + \frac{1}{3} \kappa (1 + A_\kappa \theta^2) S^3, \end{aligned} \quad (3.9)$$

where

$$\begin{aligned} m_i &\sim A_{\lambda, \kappa} \sim B_{1,2} \sim m_{\text{soft}}, \\ \mu_{1,2} &\sim m_{\text{soft}} \left(\frac{v_{\text{PQ}}}{M_{\text{Pl}}} \right)^{n_{1,2}} \quad (n_{1,2} \geq 0), \\ \kappa &\sim \left(\frac{v_{\text{PQ}}}{M_{\text{Pl}}} \right)^{n_0} \text{ or } \left(\frac{m_{\text{soft}}}{M_{\text{Pl}}} \right) \left(\frac{v_{\text{PQ}}}{M_{\text{Pl}}} \right)^{n_0} \quad (n_0 \geq 1), \end{aligned} \quad (3.10)$$

with a PQ scale given by $v_{\text{PQ}} \sim \sqrt{m_{\text{soft}} M_{\text{Pl}}}$. A simple generic feature of the PQ-NMSSM is that the singlet cubic coupling κ is always negligible,

$$\kappa \lesssim \mathcal{O}(v_{\text{PQ}}/M_{\text{Pl}}) \sim 10^{-7} - 10^{-8}, \quad (3.11)$$

while the singlet mass parameters $\mu_{1,2}$ can be either of the order of m_{soft} or negligibly small compared to m_{soft} , depending on the relative charge between S and PQ-breaking fields.

³The axion vacuum value is not determined by SUSY breaking effects, but fixed by the low energy QCD dynamics at a value solving the strong CP problem.

3.2 A minimal PQ-NMSSM

In this subsection, we present a specific model which is considered to be a minimal PQ-invariant NMSSM, and discuss the neutralino sector of the model. At high scales above v_{PQ} , but below the Planck scale M_{Pl} , the model includes the PQ-breaking superfields X and Y , as well as the NMSSM Higgs superfields S , H_u and H_d , with the following PQ-charges:

$$(S, H_u H_d, X, Y) = \left(1, -1, \frac{1}{2}, -\frac{1}{6}\right). \quad (3.12)$$

The model can include also exotic gauge-charged matter superfields Ψ_I, Ψ_I^c ($I = 1, 2$), which are vector-like under the SM gauge group, e.g. $\mathbf{5} + \bar{\mathbf{5}}$ of $\text{SU}(5)$, and carry a PQ-charge which allows renormalizable Yukawa couplings to X or Y , e.g.

$$(\Psi_1 \Psi_1^c, \Psi_2 \Psi_2^c) = \left(-\frac{1}{2}, \frac{1}{6}\right). \quad (3.13)$$

Then the most general PQ-invariant Kähler potential and superpotential are written as

$$K = \sum_i |\Phi_i|^2 + \frac{1}{M_{Pl}} (y_1 X^2 S^* + \text{h.c.}) + \dots, \\ W = \lambda S H_u H_d + \lambda' X \Psi_1 \Psi_1^c + \lambda'' Y \Psi_2 \Psi_2^c + \frac{1}{M_{Pl}} (y_2 X^2 H_u H_d + y_3 X Y^3) + \dots, \quad (3.14)$$

where the ellipses denote higher dimensional operators suppressed by higher powers of $1/M_{Pl}$. Including soft SUSY breaking terms, the scalar potential of the PQ-breaking fields takes the form

$$V = m_X^2 |X|^2 + m_Y^2 |Y|^2 + \left(\frac{y_3 A_3}{M_{Pl}} X Y^3 + \text{h.c.}\right) + \frac{y_3^2}{M_{Pl}^2} |Y|^6 + \dots. \quad (3.15)$$

Assuming $m_Y^2 < 0$ and $m_X^2 > 0$ around the renormalization point $\sim \sqrt{m_{\text{soft}} M_{Pl}}$, which can be a consequence of either a D -term induced soft SUSY breaking or the radiative correction due to a large Yukawa coupling λ'' , one finds

$$\langle X \rangle \sim \langle Y \rangle \sim \sqrt{m_{\text{soft}} M_{Pl}}, \quad \frac{F^X}{X} \sim \frac{F^Y}{Y} \sim m_{\text{soft}}, \quad (3.16)$$

assuming that

$$|m_X| \sim |m_Y| \sim |A_3| \sim m_{\text{soft}}, \quad y_3 = \mathcal{O}(1).$$

Now we can replace the PQ-breaking superfields X and Y with their vacuum expectation values while including the soft SUSY-breaking terms explicitly. Making further a field redefinition of (3.7) and a Kähler transformation of (3.8), we find that the resulting low energy effective theory takes the form

$$K_{\text{eff}} = \sum_i (1 - m_i^2 \theta^2 \bar{\theta}^2) |\Phi_i|^2, \\ W_{\text{eff}} = (\text{MSSM Yukawa and } A\text{-terms}) + \lambda(1 + A_\lambda \theta^2) S H_u H_d + \mu_1^2 (1 + B_1 \theta^2) S, \quad (3.17)$$

where the quadratic and cubic terms of S in W_{eff} are omitted since their coefficients are negligibly small:

$$\mu_2 \sim m_{\text{soft}} \left(\frac{v_{\text{PQ}}}{M_{Pl}} \right)^4, \quad \kappa \sim \left(\frac{m_{\text{soft}}}{M_{Pl}} \right) \left(\frac{v_{\text{PQ}}}{M_{Pl}} \right)^6.$$

Although the effective superpotential of this minimal PQ-NMSSM takes a simple form,⁴ the Higgs sector of the model is not distinctive as the Higgs mixing parameter $\Lambda = A_\lambda + \langle \partial_S^2 W_{\text{eff}} \rangle = A_\lambda$ is still independent from $B\mu$. On the other hand, the neutralino sector of the model is quite distinctive since $\partial_S^2 W_{\text{eff}} = 0$, and therefore the singlino gains a mass only through the mixing with other neutralinos:

$$-\mathcal{L}_{\chi_i^0} = \mu \tilde{H}_u^0 \tilde{H}_d^0 + \lambda v \cos \beta \tilde{H}_u^0 \tilde{S} + \lambda v \sin \beta \tilde{H}_d^0 \tilde{S} + \cdots, \quad (3.18)$$

where the ellipsis denotes the gaugino mass and gaugino-higgsino mixing terms. It can be shown that the lightest neutralino,

$$\chi_1^0 = N_{11} \tilde{B} + N_{12} \tilde{W}^0 + N_{13} \tilde{H}_d^0 + N_{14} \tilde{H}_u^0 + N_{15} \tilde{S},$$

has a mass lighter than $\lambda v \cos \beta$ in the limit when the mixing with gauginos is ignored [34]. To see qualitatively the properties of the lightest neutralino in the minimal PQ-NMSSM, one can take the limit of $\mu \gg \lambda v$ and the gaugino masses $M_i \gg v$. Then the neutralino mixing coefficients are found to be

$$\begin{aligned} N_{13} &= -\frac{\lambda v \cos \beta}{\mu} + \mathcal{O}((\lambda v/\mu)^2), \\ N_{14} &= -\frac{\lambda v \sin \beta}{\mu} + \mathcal{O}((\lambda v/\mu)^2), \\ N_{15} &= 1 - \frac{\lambda^2 v^2}{2\mu^2} + \mathcal{O}((\lambda v/\mu)^3), \end{aligned} \quad (3.19)$$

while the mass eigenvalue is given by

$$m_{\chi_1^0} = \frac{\lambda^2 v^2 \sin 2\beta}{\mu} (1 + \mathcal{O}((\lambda v/\mu)^2)). \quad (3.20)$$

The gaugino components in χ_1^0 are generally small because they are further suppressed by $v/M_i \ll 1$:

$$|N_{1i}| = g_i \frac{\lambda v^2 \cos 2\beta}{\sqrt{2}\mu M_i} (1 + \mathcal{O}(\lambda v/\mu)), \quad (3.21)$$

for $i = 1, 2$ with M_i being the corresponding gaugino mass.

There are important constraints on the minimal PQ-NMSSM associated with the small mass of the lightest neutralino. One is from the LEP bound on the neutralino production via the Z -boson exchange [55]:

$$\sigma(e^+e^- \rightarrow \chi_2^0 \chi_1^0) \times \text{Br}(\chi_2^0 \rightarrow q\bar{q}\chi_1^0) \lesssim 100 \text{ fb}, \quad (3.22)$$

⁴Note that W_{eff} of the minimal PQ-NMSSM is the same as that of the nMSSM, which has been proposed in refs. [49–51] in a different context. See also refs. [52–54].

which applies for $m_{\chi_2^0} + m_{\chi_1^0} < 208 \text{ GeV}$ and $m_{\chi_1^0} > 60 \text{ GeV}$. This puts an upper bound on the Z -boson coupling to $\chi_2^0 \chi_1^0$. In addition, the global fit analysis excludes an invisible decay of the SM-like Higgs boson with a branching ratio greater than 0.38 at 95% confidence level, if one allows its couplings to the SM particles to deviate from the SM values [56, 57]:

$$\text{Br}(h \rightarrow \chi_1^0 \chi_1^0) < 0.38. \quad (3.23)$$

In the NMSSM, the Higgs coupling for this process is given by

$$y_{h\chi_1^0\chi_1^0} = \frac{\sqrt{2}\lambda^2 v \sin 2\beta}{\mu} (1 + \mathcal{O}(\lambda v/\mu)), \quad (3.24)$$

which can have a sizable value in the limit of low $\tan\beta$, large λ , and light μ . To avoid a dangerous Higgs invisible decay when the Higgs coupling $y_{h\chi_1^0\chi_1^0}$ is sizable, one needs $2m_{\chi_1^0} > m_h$ so that the process is kinematically forbidden. On the other hand, if $2m_{\chi_1^0} < m_h$, one needs to adjust the model to suppress $y_{h\chi_1^0\chi_1^0}$. However, with small $y_{h\chi_1^0\chi_1^0}$, it is difficult to have a sizable NMSSM contribution to the tree level mass of the SM-like Higgs boson, which is the feature that we like to keep to avoid too severe fine-tuning of the model. Note that, since we are assuming $m_0 \lesssim 120 \text{ GeV}$, a sizable NMSSM contribution is required to get $m_h \simeq 125 \text{ GeV}$. Actually, as we shall see in the next section, this makes it difficult to suppress the branching fraction for the Higgs invisible decay below 0.38 in most of the parameter space of our interest once the mode is kinematically open. We therefore require $2m_{\chi_1^0} > m_h$ to prohibit the decay process $h \rightarrow \chi_1^0 \chi_1^0$.

4 Higgs phenomenology of the PQ-NMSSM

The SM-like Higgs boson observed at the LHC can be accommodated in the NMSSM while satisfying the constraints on scalar mixing discussed in section 2. We will first examine how large the mixing between the SM-like Higgs boson and the singlet-like Higgs boson is allowed in the general NMSSM, and then move on to the minimal PQ-NMSSM where the mixing is further constrained due to a light singlino-like neutralino. As we will see, a SM-like Higgs boson with sizable singlet component can be compatible with all the LHC and LEP results available at present. In such case, the singlet scalar is expected to be around the weak scale since otherwise a sizable singlet-doublet mixing would make it difficult to explain the observed SM-like Higgs boson mass $m_h \simeq 125 \text{ GeV}$. The singlet-like Higgs boson s can be lighter or heavier than the SM-like Higgs boson h . For the former case, we will focus on the possibility that the 2σ excess of the LEP $Zb\bar{b}$ events at $m_{b\bar{b}} \simeq 98 \text{ GeV}$ is explained by s with $m_s \simeq 98 \text{ GeV}$.

Let us briefly explain how we explore the effect of scalar mixing. The relations (2.12) and (2.20) will be used to express $\{\lambda, \mu, m_0\}$ in terms of $\{\theta_i, \tan\beta, m_s, m_H\}$ with $m_h \simeq 125 \text{ GeV}$. We also require that the scalar mixing angles obey the relation (2.16) in order for the Higgs signal rate for the WW/ZZ channel to be close to the SM prediction. Then, taking some benchmark values of the Higgs boson masses m_s and m_H , one can see how the model parameters $\{\lambda, \mu, m_0\}$ change on the parameter plane $(\theta_2, \tan\beta)$ for a given value of θ_3 . In other words, it is possible to figure out which region in the $(\theta_2, \tan\beta)$ space is

allowed by the constraints on $\{\lambda, \mu, m_0\}$. Over the allowed region, we will examine the signal rates of the SM-like Higgs boson for the $f\bar{f}$ ($f = b, \tau$) and $\gamma\gamma$ channel, and also the properties of the singlet-like Higgs boson.

4.1 Singlet-like Higgs boson at 98 GeV

If the singlet-like Higgs boson s is lighter than 114 GeV, scalar mixing is constrained not only by the LHC results, but also by the LEP search of the Higgs boson. In regard to this possibility, a particularly interesting LEP result is the 2σ excess of $Zb\bar{b}$ around $m_{b\bar{b}} \simeq 98$ GeV [58, 59]. This may indicate a light singlet-like Higgs boson with $m_s \simeq 98$ GeV, which can have a sizable coupling to the Z boson through the mixing with the SM-like Higgs boson:

$$C_V^s \approx \theta_2. \quad (4.1)$$

Let $R_s^{Zb\bar{b}}$ be the signal strength for $e^+e^- \rightarrow Zs \rightarrow Zb\bar{b}$ normalized by the signal rate for the SM Higgs boson having the same mass. In the case that s decays dominantly into a bottom quark pair, which is indeed the case for the parameter space of our interest, we have $R_s^{Zb\bar{b}} \simeq |C_V^s|^2$, and thus the LEP excess can be explained if

$$0.1 \lesssim \theta_2^2 \lesssim 0.25, \quad m_s \simeq 98 \text{ GeV}. \quad (4.2)$$

Let us examine how $\{\lambda, \mu, m_0\}$ change on the $(\theta_2, \tan\beta)$ plane for the case that s explains the LEP excess around 98 GeV, and h has the properties observed at the LHC. Imposing the condition (2.16) for $R_h^{VV} \approx 1$, together with $m_s = 98$ GeV and $m_H^2 \gg v^2$, one finds that λ and μ can be determined in terms of θ_2 , θ_3 , and $\tan\beta$ according to

$$\begin{aligned} \lambda^2 \approx & 0.27 - 0.10 \tan\beta \theta_2 \theta_3 \\ & + 1.11 \left(\frac{m_H}{350 \text{ GeV}} \right)^2 \left(1 - \frac{m_h^2}{m_H^2} \right) \left(\theta_2^2 - \frac{1.4 \tan^2\beta + 1.7}{\tan\beta} \theta_2 \theta_3 \right), \end{aligned} \quad (4.3)$$

$$\mu \approx 100 \text{ GeV} \left(-0.43 \theta_2 + 16.2 \left(\frac{m_H}{350 \text{ GeV}} \right)^2 \frac{\theta_3}{\tan\beta} \right) \left(\frac{\lambda}{0.4} \right)^{-1}, \quad (4.4)$$

in the expansion in powers of θ_i . Here we have taken into account that the charged Higgs scalar, whose mass is similar to m_H , should be heavier than about 350 GeV to satisfy the $b \rightarrow s\gamma$ constraint, barring cancellation with other superparticle contributions [60]. Similarly, one also finds

$$\begin{aligned} \frac{m_0^2}{m_h^2} \approx & 1 - 0.4 \left(\theta_2^2 - \frac{2}{\tan\beta} \theta_2 \theta_3 \right) \\ & - \frac{11}{\tan^2\beta} \left(\frac{m_H}{350 \text{ GeV}} \right)^2 \left(1 - \frac{m_h^2}{m_H^2} \right) \left(\theta_2^2 - \frac{1.4 \tan^2\beta + 1.7}{\tan\beta} \theta_2 \theta_3 \right). \end{aligned} \quad (4.5)$$

Although a naive approximation, the above relations help us to qualitatively understand the effect of scalar mixing for a given value of m_H . For $105 \text{ GeV} \lesssim m_0 \lesssim 120 \text{ GeV}$, the last term in (4.5) cannot be large, constraining Higgs mixing to be

$$0 \lesssim \delta_\theta \equiv \theta_2^2 - \frac{1.4 \tan^2\beta + 1.7}{\tan\beta} \theta_2 \theta_3 \lesssim 0.04 \left(\frac{m_H}{350 \text{ GeV}} \right)^{-2} \tan^2\beta. \quad (4.6)$$

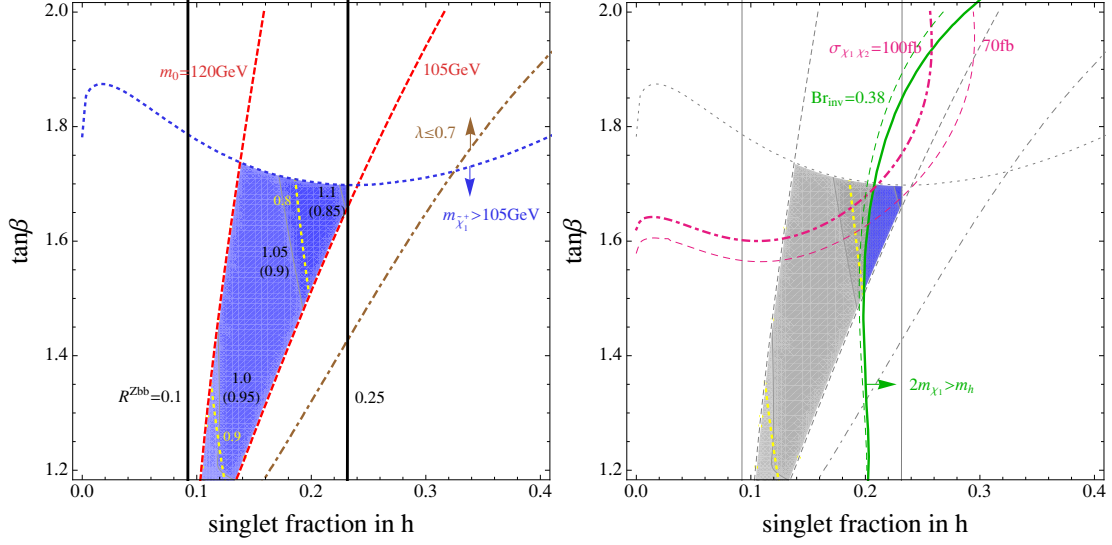


Figure 1. Singlet fraction ($= c_{\theta_1}^2 s_{\theta_2}^2$) of the SM-like Higgs boson and $\tan\beta$ consistent with (a) $m_h \simeq 125$ GeV with $105 \text{ GeV} \leq m_0 \leq 120 \text{ GeV}$ (the region between the two dashed red lines), (b) $R_h^{VV} \approx 1$, (c) $\lambda \leq 0.7$ (the region above the dot-dashed brown line), and (d) the LEP bound on the chargino mass (the region below the dotted blue line). We also impose $0.1 \leq R_s^{Zb\bar{b}} \leq 0.25$ to fit the LEP excess of $Zb\bar{b}$ at $m_{b\bar{b}} \simeq 98$ GeV (the region between the two vertical black lines). The left panel is the result for the general NMSSM, where we have taken $m_s = 98$ GeV, $m_H = 350$ GeV, $|\theta_3| = 0.1$ with $\theta_2\theta_3 > 0$, and $2M_1 = M_2 = 300$ GeV. The contours of the Higgs signal strengths $R_h^{\gamma\gamma}/R_h^{VV}$ (thin gray line) and $R_h^{b\bar{b}}/R_h^{VV}$ (dashed yellow line) are depicted, where the number in the bracket is the di-photon rate for the opposite sign of θ_2 . The right panel shows a viable region of the minimal PQ-NMSSM, where the mixing is further constrained by $2m_{\chi_1^0} > m_h$ (the right side of the solid green line) and $\sigma(e^+e^- \rightarrow \chi_2^0\chi_1^0) < 100$ fb (the region below the thick-dot-dashed magenta line). Note that the branching fraction of the Higgs invisible decay, if kinematically open, is smaller than 0.38 only in the narrow region between the dashed and solid green lines.

Combining this with the relation (4.4), the chargino mass bound $|\mu| \gtrsim 100$ GeV leads to an upper bound on $\tan\beta$,

$$\tan\beta \lesssim 1.9 \left(\frac{m_H}{350 \text{ GeV}} \right) \left(\frac{\theta_2^2}{0.25} \right)^{1/4} \left(\frac{\lambda}{0.4} \right)^{-1/2}. \quad (4.7)$$

One can see that the perturbativity bound $\lambda \lesssim 0.7$ can be easily satisfied for the scalar mixing angles and $\tan\beta$ satisfying (4.6) and (4.7). Note also that, because of the constraint (4.6), *positive* $\theta_2\theta_3$ is favored for the mixing angle θ_2 to be sizable.

The left plot of figure 1 illustrates the range of the singlet fraction ($= c_{\theta_1}^2 s_{\theta_2}^2$) of the SM-like Higgs boson and $\tan\beta$ for which the LEP excess of $Zb\bar{b}$ at $m_{b\bar{b}} \simeq 98$ GeV is explained by a singlet-like scalar s , while satisfying the perturbativity bound on λ , the LEP bound on chargino mass, and $105 \text{ GeV} \leq m_0 \leq 120 \text{ GeV}$. Here we have imposed the relation (2.16) to have $R_h^{VV} \approx 1$,⁵ and used $|\theta_3| = 0.1$, $m_H = 350$ GeV, and the gaugino masses $2M_1 = M_2 = 300$ GeV for the purpose of illustration. The LEP bound on the

⁵One may consider a case where the Higgs signal rate into WW/ZZ deviates from the SM value by an

chargino mass puts a lower bound on $|\mu|$, which can be relaxed if the wino mass M_2 is around a few hundred GeV and $\mu M_2 < 0$. One can see that the allowed blue-shaded region is determined mainly by the constraints associated with μ and m_0 , whose characteristic features can be understood by the relation (4.4) and (4.5). Note that $\tan\beta$ is bounded from above by the constraint on $|\mu|$ according to (4.7), while the constraint on m_0 explains the allowed range of θ_2^2 for a given value of θ_3 and $\tan\beta$ through the relation (4.6). The allowed region becomes smaller if one increases the wino mass or changes its sign, because then the lower bound on μ from the chargino mass bound is strengthened. We also present in figure 1 the contours of R_h^{ii}/R_h^{VV} for the $b\bar{b}$ ($\tau\bar{\tau}$) and di-photon channels. For the scalar mixing giving $R_h^{VV} \approx 1$, the $b\bar{b}$ ($\tau\bar{\tau}$) signal rate is always below the WW/ZZ signal rate. On the other hand, the di-photon signal rate can be either below or above the WW/ZZ signal rate depending on the sign of the singlet-doublet mixing angle θ_2 , because the Higgs coupling to photons receives a sizable contribution from the charged-higgsino loop when the higgsino mass is around the weak scale.⁶

Now we impose the additional constraints that are particularly relevant for the minimal PQ-NMSSM which predicts a light singlino-like neutralino:

$$2m_{\chi_1^0} > m_h, \quad \sigma(e^+e^- \rightarrow \chi_2^0\chi_1^0) \lesssim 100 \text{ fb}. \quad (4.9)$$

As shown in the right plot of figure 1, only a small region remains viable, in which R_h^{bb}/R_h^{VV} is about 0.8, and $R_h^{\gamma\gamma}/R_h^{VV}$ deviates from one by about ± 0.1 depending on the sign of θ_2 . On the other hand, notice that the branching fraction for the Higgs invisible decay is larger than 0.38 in most of the parameter space where the decay mode is kinematically allowed, except in the narrow region between the dashed and solid green lines where there is a large kinematic suppression. A notable feature of the minimal PQ-NMSSM is that a phenomenologically viable parameter region points toward stops around or below 1 TeV and light higgsinos around the weak scale.

Figure 2 shows how the allowed region changes with m_H and θ_3 . The left panel is obtained by taking a heavier m_H compared to figure 1, while the right panel is the result for a larger value of $|\theta_3|$. The relation (4.7) indicates that the upper bound on $\tan\beta$ increases as H becomes heavier, while the relation (4.6) explains why the shaded region in the figure is reduced for heavier m_H and why it moves to the right when one takes larger $|\theta_3|$. On the other hand, for a given $\tan\beta$, μ becomes large if one raises m_H or θ_3 as can be seen in (4.4). A large μ makes it more difficult to satisfy the condition $2m_{\chi_1^0} > m_h$, so a

amount δR_h^{VV} due to scalar mixing. Then, the relation (2.16) should be modified as

$$\theta_1 \approx \frac{\tan\beta}{1.4\tan^2\beta + 1.7} \left(\theta_2^2 + \delta R_h^{VV} \right), \quad (4.8)$$

and the mixing effects can be examined by taking the replacement $\theta_2^2 \rightarrow \theta_2^2 + \delta R_h^{VV}$ in the relations (4.3) and (4.5). As a result, the region consistent with $105 \text{ GeV} \leq m_0 \leq 120 \text{ GeV}$ will move horizontally to the left (right) in figure 1 if δR_h^{VV} is positive (negative).

⁶The Higgs coupling to photons receives a loop contribution also from the $h\tilde{H}^+\tilde{W}^-$ interaction, which becomes important when both higgsinos and winos have masses not much above the weak scale. Such an effect has been included in our analysis.

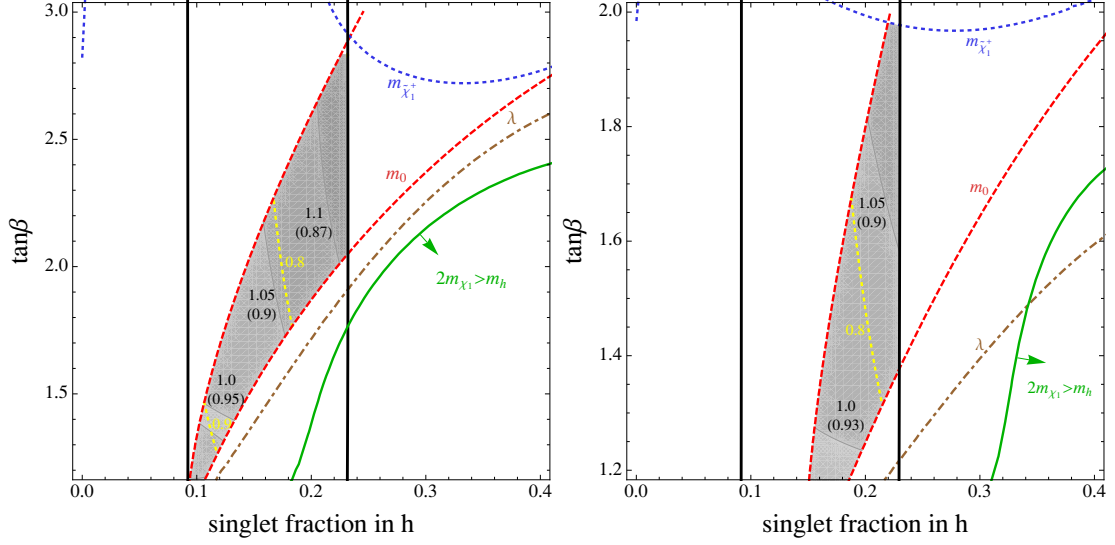


Figure 2. Higgs mixing consistent with various constraints, where the notations are the same as in figure 1. The left panel shows an allowed region for the general NMSSM with $m_H = 500$ GeV and $|\theta_3| = 0.1$, while the right panel is for $m_H = 350$ GeV and $|\theta_3| = 0.12$. In both cases, we have taken $m_s = 98$ GeV, $\theta_2\theta_3 > 0$ and $2M_1 = M_2 = 300$ GeV. For heavier m_H (500 GeV) or larger $|\theta_3|$ (0.12), there is no viable parameter region satisfying $2m_{\chi_1^0} > m_h$ or $\text{Br}(h \rightarrow \chi_1^0\chi_1^0) < 0.38$ in the minimal PQ-NMSSM.

phenomenologically viable region of the minimal PQ-NMSSM gets smaller, or disappears, as m_H or θ_3 increases.

We close this subsection by pointing out that the minimal PQ-NMSSM requires stops around or below 1 TeV, and higgsinos around the weak scale. If m_0 is larger than about 110 GeV, it is difficult to have $2m_{\chi_1^0} > m_h$, which would be necessary to forbid $h \rightarrow \chi_1^0\chi_1^0$. This means that h can be identified as the SM-like Higgs boson observed at the LHC only when stops are not significantly heavier than 1 TeV. In addition, combined with $m_0 \gtrsim 105$ GeV, the requirement $2m_{\chi_1^0} > m_h$ constrains μ to be around the weak scale. As we will see in the next subsection, these features hold also for the case that s is heavier than h .

4.2 Singlet-like Higgs boson above 125 GeV

Let us move to the case where the singlet-like Higgs boson s is heavier than the SM-like Higgs boson h . One of the main differences from the opposite case with $m_s < m_h$ is that the h - s mixing always decreases m_h . Thus we need $\lambda > m_Z/v$ and low $\tan\beta$ in order to arrange $m_h \simeq 125$ GeV in the presence of scalar mixing, unless m_0 is larger than 125 GeV. It is clear that the singlet-doublet mixing angle θ_2 can be sizable if s is not much heavier than h . As in the previous case, the effect of scalar mixing can be understood qualitatively by using the approximated relations (4.3) and (4.5) after multiplying the second term with $(m_s^2 - m_h^2)/((98\text{GeV})^2 - m_h^2)$, and the relation (4.4) after multiplying the first term with $m_s^2/(98\text{GeV})^2$. Then it follows that a viable region for the general NMSSM with $m_s > m_h$ appears at lower $\tan\beta$ compared to the case with $m_s = 98$ GeV. Hence, it

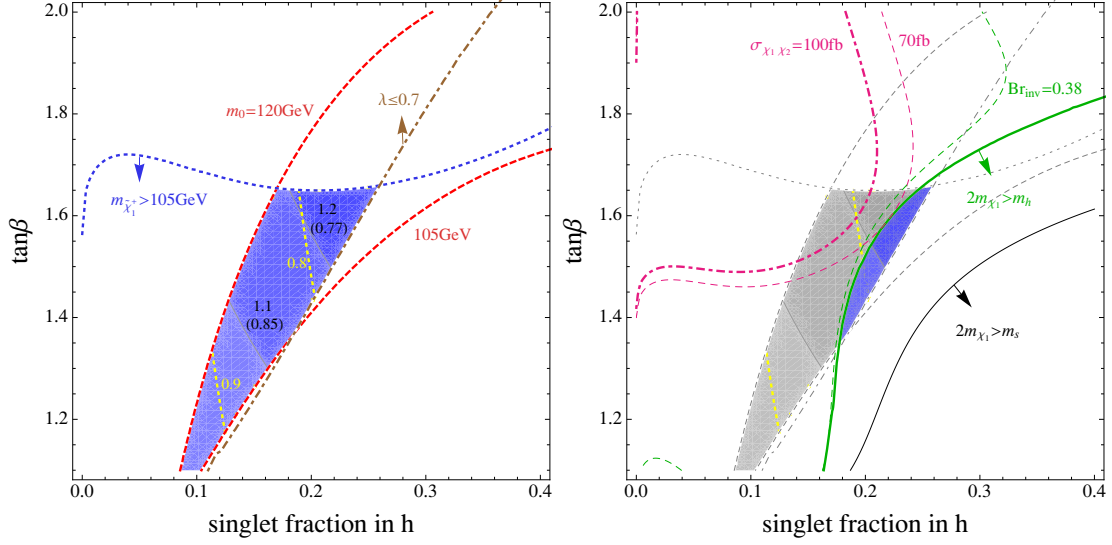


Figure 3. Higgs mixing consistent with various constraints, where the notations are the same as in figure 1. The left panel shows a viable region in the general NMSSM with $m_s = 150$ GeV, $m_H = 350$ GeV, and $|\theta_3| = 0.1$ with $\theta_2\theta_3 > 0$, while the right panel is the result for the minimal PQ-NMSSM. Here the gauginos are assumed to be much heavier than the weak scale.

becomes relatively easy to satisfy the condition $2m_{\chi_1^0} > m_h$ in the minimal PQ-NMSSM where $m_{\chi_1^0}$ is proportional to $\sin 2\beta$.

In the left panel of figure 3, we show the region of $(c_{\theta_1}^2 s_{\theta_2}^2, \tan \beta)$ compatible with the constraints on $\{\lambda, \mu, m_0\}$ for the general NMSSM with $m_s = 150$ GeV, $m_H = 350$ GeV, and $|\theta_3| = 0.1$. Here, for simplicity, we have assumed that the gauginos are much heavier than the weak scale so they are decoupled well from the singlino-like neutralino. The figure shows that, for $m_s > m_h$, the perturbativity bound on λ is as important as the other constraints. However a sizable singlet fraction of h is still allowed. The right panel shows the result for the minimal PQ-NMSSM, which is further constrained by the bound on the Higgs invisible decay rate and the LEP bound on the neutralino production rate. Again, there is an allowed region with sizable θ_2 , in which R_h^{bb}/R_h^{VV} is around 0.8 for the scalar mixing consistent with $R_h^{VV} \approx 1$, while $R_h^{\gamma\gamma}/R_h^{VV}$ deviates from one by about ± 0.2 depending on the sign of θ_2 .

On the other hand, if the singlet-like Higgs becomes heavier, only a smaller value of θ_2 will be allowed. For instance, the singlet fraction of h should be less than about 0.1 for $m_s > 250$ GeV and $m_H = 350$ GeV. Figure 4 illustrates how the allowed region of $(c_{\theta_1}^2 s_{\theta_2}^2, \tan \beta)$ varies when m_H becomes heavier (left panel) or θ_3 becomes larger (right panel). We can see that the allowed region varies in the same way as explained in the previous subsection, so that the phenomenologically viable region of the minimal PQ-NMSSM becomes smaller or vanishes as m_H or θ_3 gets larger.

It is worth noting that, if s is heavy enough, it can dominantly decay into a pair of the lightest neutralino. For instance, for the minimal PQ-NMSSM depicted in the right panel of figure 3, the region above the thin black line gives $m_s > 2m_{\chi_1^0}$, for which the invisible decay $s \rightarrow \chi_1^0 \chi_1^0$ is open. This region covers all the parameter space satisfying the

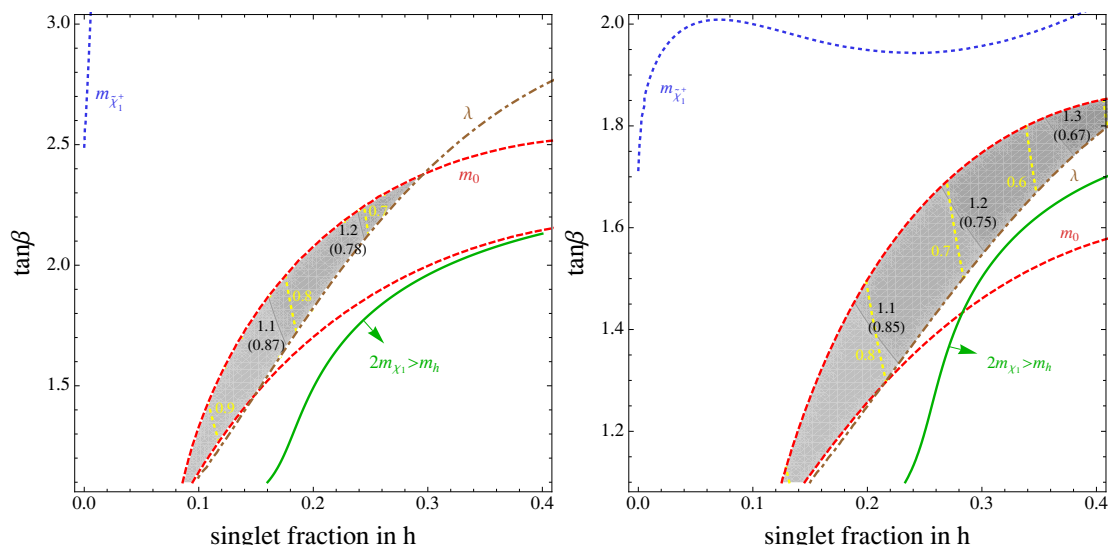


Figure 4. Higgs mixing consistent with various constraints, where the notations are the same as in figure 1. The left panel shows an allowed region in the general NMSSM with $m_H = 500$ GeV and $|\theta_3| = 0.1$, while the right panel is for $m_H = 350$ GeV and $|\theta_3| = 0.12$. In both cases, we have taken $m_s = 150$ GeV, $\theta_2\theta_3 > 0$, and assumed the heavy gauginos. The minimal PQ-NMSSM with heavier m_H (500 GeV) or larger $|\theta_3|$ (0.12) does not have a viable parameter region satisfying $2m_{\tilde{\chi}_1^0} > m_h$ or $\text{Br}(h \rightarrow \chi_1^0\chi_1^0) < 0.38$.

phenomenological constraints discussed here. We find that the branching fraction of the invisible decay $s \rightarrow \chi_1^0\chi_1^0$ is about 0.7 – 0.8 over the viable region, which would make it difficult to discover s at collider experiments.

5 Conclusions

It is an interesting possibility that there exist additional light Higgs bosons near the weak scale other than the SM-like Higgs boson recently discovered at the LHC. The next-to-minimal supersymmetric standard model (NMSSM) is a natural place to realize this scenario because it can accommodate a SM-like 125 GeV Higgs boson with less fine-tuning, and additional scalar bosons can have a light mass without causing a further fine-tuning. The axion solution to the strong CP problem can be easily incorporated in the NMSSM. Then an appealing scenario is that an intermediate PQ scale $v_{\text{PQ}} \sim \sqrt{m_{\text{soft}}\overline{M}_{Pl}}$ emerges through an interplay between SUSY breaking effect and Planck-scale suppressed effect, and low energy mass parameters of $\mathcal{O}(m_{\text{soft}})$ in the effective superpotential are induced by a spontaneous breakdown of the PQ symmetry.

We have examined the Higgs phenomenology in such a PQ-invariant NMSSM while focusing on the phenomenological consequences of scalar mixing. The observed mass and signal rates of the SM-like Higgs boson and the LEP bound on the chargino mass provide important constraints on scalar mixing. We imposed also the perturbativity bound on the singlet Yukawa coupling λ , and assumed stop masses between 600 GeV and a few TeV. In addition to these constraints, the minimal PQ-NMSSM is further constrained due to

the presence of a light singlino-like neutralino in the model, most notably by the invisible decay of the SM-like Higgs boson into a neutralino pair.

The singlet-doublet mixing can also affect the mass of the SM-like Higgs boson h . For $m_s < m_h$, we have explored the possibility that the 2σ excess of the LEP $Zb\bar{b}$ events at $m_{b\bar{b}} \simeq 98 \text{ GeV}$ is explained by a singlet-like Higgs boson s with $m_s \simeq 98 \text{ GeV}$ within the framework of the minimal PQ-NMSSM. Interestingly enough, this requires a low $\tan\beta$ smaller than about 2, together with stops around or below 1 TeV, and light doublet-higgsinos around the weak scale. On the other hand, for the case with $m_s > m_h$, we found that s dominantly decays into neutralinos in most of the phenomenologically viable parameter region. In both cases, the signal rate of the SM-like Higgs boson decaying into $b\bar{b}$ or $\tau\bar{\tau}$ is reduced by 10–20% compared to the signal rate into WW/ZZ , while the signal rate in the di-photon channel is reduced or enhanced by a similar amount depending on the sign of the singlet-doublet mixing angle. Here scalar mixing has been constrained by the observational requirement that the Higgs signal rate into WW/ZZ be close to the SM value.

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A Scalar masses and mixings in the NMSSM

For the CP-even neutral Higgs bosons defined as

$$\begin{aligned}\hat{h} &= \sqrt{2} \left((\text{Re}H_d^0 - v \cos\beta) \cos\beta + (\text{Re}H_u^0 - v \sin\beta) \sin\beta \right), \\ \hat{H} &= \sqrt{2} \left((\text{Re}H_d^0 - v \cos\beta) \sin\beta - (\text{Re}H_u^0 - v \sin\beta) \cos\beta \right), \\ \hat{s} &= \sqrt{2} (\text{Re}S - \langle S \rangle),\end{aligned}\tag{A.1}$$

the scalar mass matrix is given by

$$\hat{M}^2 = \begin{pmatrix} m_h^2 & \frac{1}{2}(m_Z^2 - \lambda^2 v^2) \sin 4\beta & \lambda v(2\mu - \Lambda \sin 2\beta) \\ \frac{1}{2}(m_Z^2 - \lambda^2 v^2) \sin 4\beta & m_{\hat{H}}^2 & \lambda v \Lambda \cos 2\beta \\ \lambda v(2\mu - \Lambda \sin 2\beta) & \lambda v \Lambda \cos 2\beta & m_{\hat{s}}^2 \end{pmatrix}, \tag{A.2}$$

where $\langle |H_u^0| \rangle = v \sin\beta$ and $\langle |H_d^0| \rangle = v \cos\beta$, and

$$m_h^2 = m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta + \frac{3m_t^4}{4\pi^2 v^2} \ln \left(\frac{m_t^2}{m_t^2} \right) + \frac{3m_t^4}{4\pi^2 v^2} \left(X_t^2 - \frac{1}{12} X_t^4 \right), \tag{A.3}$$

$$m_{\hat{H}}^2 = \frac{2B\mu}{\sin 2\beta} - (\lambda^2 v^2 - m_Z^2) \sin^2 2\beta, \tag{A.4}$$

for the stop mixing parameter $X_t = (A_t - \mu \cot \beta)/m_{\tilde{t}}$, and $\Lambda = A_\lambda + \langle \partial_S^2 f \rangle$. Here m_s^2 depends on the superpotential f for the singlet S , and the soft scalar mass term for it. The mass matrix is diagonalized by an orthogonal matrix,

$$U = \begin{pmatrix} c_{\theta_1} c_{\theta_2} & -s_{\theta_1} & -c_{\theta_1} s_{\theta_2} \\ s_{\theta_1} c_{\theta_2} c_{\theta_3} - s_{\theta_2} s_{\theta_3} & c_{\theta_1} c_{\theta_3} & -c_{\theta_2} s_{\theta_3} - s_{\theta_1} s_{\theta_2} c_{\theta_3} \\ s_{\theta_1} c_{\theta_2} s_{\theta_3} + s_{\theta_2} c_{\theta_3} & c_{\theta_1} s_{\theta_3} & c_{\theta_2} c_{\theta_3} - s_{\theta_1} s_{\theta_2} s_{\theta_3} \end{pmatrix}, \quad (\text{A.5})$$

where $c_\theta = \cos \theta$ and $s_\theta = \sin \theta$ for θ_i in the range between $-\pi/2$ and $\pi/2$. Assuming that the mixing angles are small, the mass eigenstate

$$h = c_{\theta_1} c_{\theta_2} \hat{h} - s_{\theta_1} \hat{H} - c_{\theta_1} s_{\theta_2} \hat{s}$$

can be identified as the observed SM-like Higgs boson with a mass $m_h \simeq 125$ GeV.

The Lagrangian parameters are related to the Higgs mass eigenvalues and mixing angles as follows:

$$\begin{aligned} \lambda^2 v^2 &= m_Z^2 + \frac{1}{\sin 4\beta} \left((m_H^2 - m_s^2) s_{\theta_2} s_{2\theta_3} + 2(m_h^2 - m_H^2 c_{\theta_3}^2 - m_s^2 s_{\theta_3}^2) s_{\theta_1} c_{\theta_2} \right) c_{\theta_1}, \\ \lambda v \mu &= -\frac{1}{4} m_h^2 c_{\theta_1}^2 s_{2\theta_2} - \frac{1}{4} (m_H^2 - m_s^2) s_{\theta_1} c_{2\theta_2} s_{2\theta_3} \\ &\quad + \frac{1}{4} \left((m_H^2 - m_s^2 s_{\theta_1}^2) s_{\theta_3}^2 - (m_H^2 s_{\theta_1}^2 - m_s^2) c_{\theta_3}^2 \right) s_{2\theta_2} \\ &\quad - \frac{\tan 2\beta}{4} \left((m_H^2 - m_s^2) c_{\theta_2} s_{2\theta_3} - 2(m_h^2 - m_H^2 c_{\theta_3}^2 - m_s^2 s_{\theta_3}^2) s_{\theta_1} s_{\theta_2} \right) c_{\theta_1}, \\ \lambda v \Lambda &= -\frac{1}{2 \cos 2\beta} \left((m_H^2 - m_s^2) c_{\theta_2} s_{2\theta_3} - 2(m_h^2 - m_H^2 c_{\theta_3}^2 - m_s^2 s_{\theta_3}^2) s_{\theta_1} s_{\theta_2} \right) c_{\theta_1}, \end{aligned} \quad (\text{A.6})$$

and

$$\begin{aligned} m_h^2 &= m_h^2 c_{\theta_1}^2 c_{\theta_2}^2 + m_H^2 (s_{\theta_1} c_{\theta_2} c_{\theta_3} - s_{\theta_2} s_{\theta_3})^2 + m_s^2 (s_{\theta_2} c_{\theta_3} + s_{\theta_1} c_{\theta_2} s_{\theta_3})^2, \\ m_H^2 &= m_h^2 s_{\theta_1}^2 + m_H^2 c_{\theta_1}^2 c_{\theta_3}^2 + m_s^2 c_{\theta_1}^2 s_{\theta_3}^2, \\ m_s^2 &= m_h^2 c_{\theta_1}^2 s_{\theta_2}^2 + m_H^2 (s_{\theta_1} s_{\theta_2} c_{\theta_3} + c_{\theta_2} s_{\theta_3})^2 + m_s^2 (c_{\theta_2} c_{\theta_3} - s_{\theta_1} s_{\theta_2} s_{\theta_3})^2. \end{aligned} \quad (\text{A.7})$$

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